

Classwork example - solution -

(1)

$$(a) \Delta t = \frac{T}{2} = \pi \sqrt{\frac{a_H^3}{\mu_S}}$$

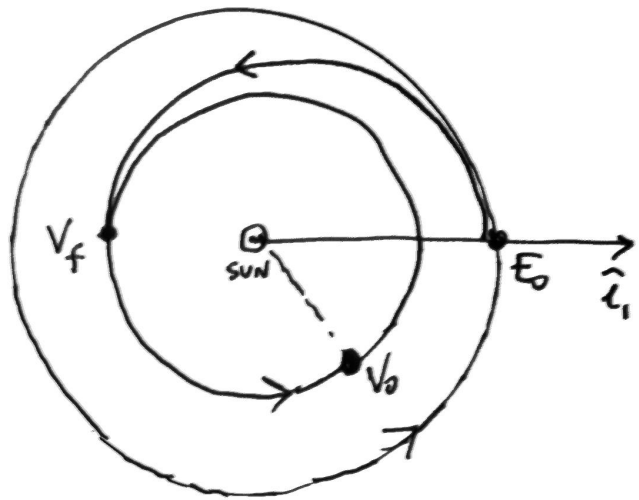
$$a_H = \frac{r_V + r_E}{2} = 128850000 \text{ km}$$

$$e_H = \frac{r_E - r_V}{r_E + r_V} = 0.160$$

$$\rightarrow \Delta t = 12613658 \text{ sec} \\ = 146.99 \text{ days}$$

$$\rightarrow \phi_{Vf} = \pi = \phi_{V0} + \sqrt{\frac{\mu_S}{r_V^3}} \Delta t$$

$\uparrow$  Angle of Venus at  $t_f$        $\uparrow$  Angle of Venus at  $t_0$



$E_{0/f}$  = Earth at initial (0) time and final (f) time

$V_{0/f}$  = Venus at initial (0) time and final (f) time

$\phi_{v/c}$  = angle taken counterclockwise from  $\hat{z}_1$

$$\rightarrow \phi_{V0} = \pi \left( 1 - \sqrt{\frac{a_H^3}{r_V^3}} \right) = -0.941 = -53.9 \text{ deg}$$

$$(b) v_{r,f}^{(H)} = 0 \quad v_{\theta,f}^{(H)} = \sqrt{\frac{\mu_S}{p_H}} (1 + e_H) = 37.722 \frac{\text{km}}{\text{sec}}$$

$$p_H = a_H (1 - e_H^2)$$

$$\underline{v_{\infty}^-} = \underline{v}^{(H)} - \underline{v}_V^{(H)} \rightarrow \underline{v_{\infty}^-} = v_{\infty} \hat{\theta} = \left( v_{\theta,f}^{(H)} - \sqrt{\frac{\mu_S}{r_V}} \right) \hat{\theta}$$

$\uparrow$                        $\uparrow$   
 Holic.                  Holic.  
 velocity of          velocity  
 spacecraft          of Venus

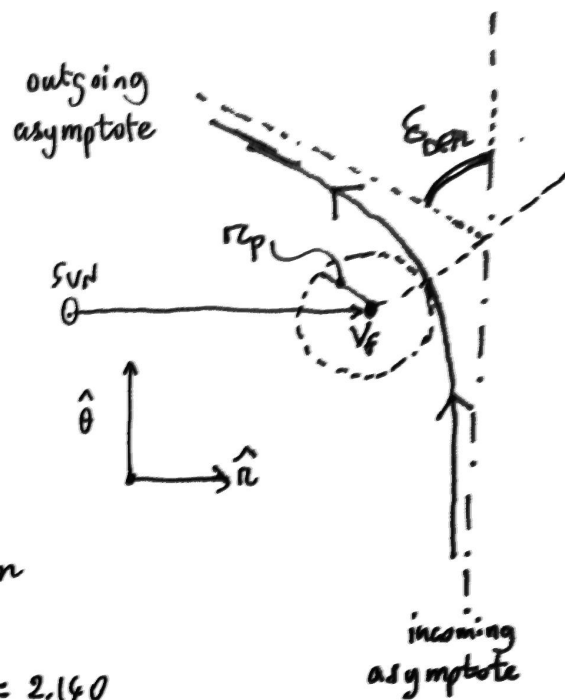
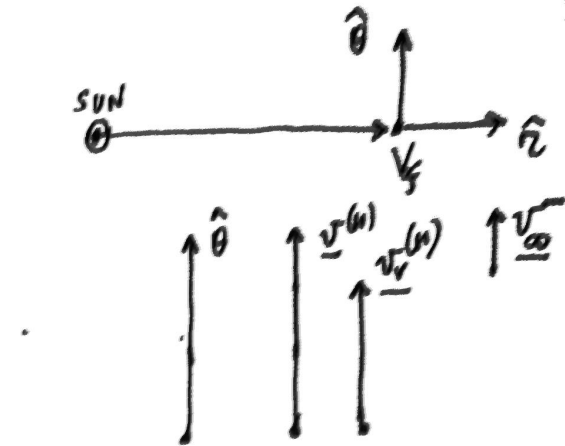
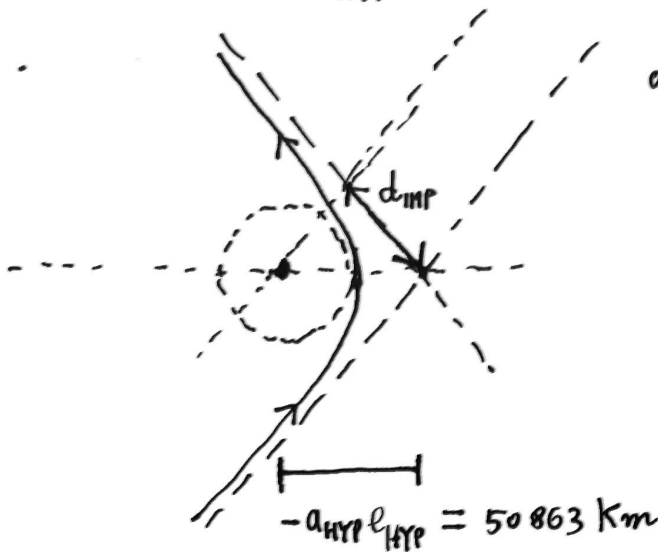
$$= 2.702 \frac{\text{km}}{\text{sec}}$$

From  $v_{\infty}$  one obtains

$$a_{HYPER} = -\frac{v_{\infty}^2}{\mu_V} = -44611 \text{ km}$$

(c)  $r_p = a_{HYPER} (1 - e_{HYPER}) = R_V + 200 \text{ km}$

$$\rightarrow e_{HYPER} = 1 - \frac{r_p}{a_{HYPER}} = 1.140$$



$$\delta_{DEFL} = 2 \theta_q^{lim} - \pi = 122.6 \text{ deg} = 2.140$$

$$d_{IMP} = -a_{HYPER} \sqrt{e_{HYPER}^2 - 1} = 24432 \text{ km}$$

(d) After flyby

$$\underline{v}^{(H)} = \underline{v}_{\infty}^+ + \underline{v}_V^{(H)} \quad \text{where} \quad \begin{cases} v_{\infty, r}^+ = -v_{\infty} \sin \delta_{DEFL} = -2.277 \frac{\text{km}}{\text{sec}} \\ v_{\infty, \theta}^+ = v_{\infty} \cos \delta_{DEFL} = -1.455 \frac{\text{km}}{\text{sec}} \end{cases}$$

In the end, after flyby the spacecraft has the following components of heliocentric velocity:

$$\begin{cases} v_r = v_{\infty, r}^+ + v_{v, r}^{(H)} = -2.777 \frac{\text{km}}{\text{sec}} \\ v_\theta = v_{\infty, \theta}^+ + v_{v, \theta}^{(H)} = 33.565 \frac{\text{km}}{\text{sec}} \end{cases}$$

and the velocity magnitude is  $v_* = 33.642 \frac{\text{km}}{\text{sec}}$   
 ( $< 37.722 \frac{\text{km}}{\text{sec}}$ , velocity prior to flyby).

(2)

(a)  $T = 2\pi \sqrt{\frac{a^3}{\mu}} \rightarrow a = 7136.635 \text{ km}$   
 $= R_E + 758.5 \text{ km}$

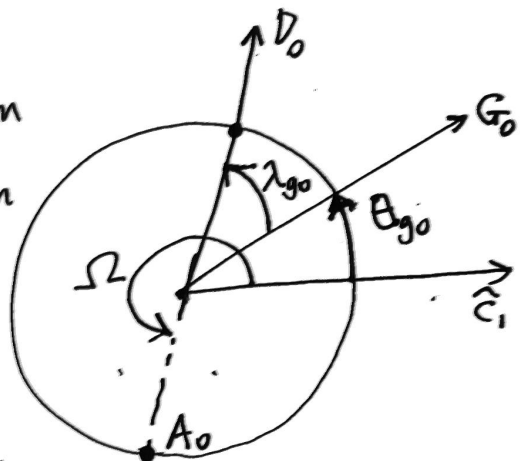
Due to  $J_2$ , the sunynchronous condition allows finding  $i$ :

$R = a$  (circular orbit)

$$i = \arccos \left[ -\frac{2\pi}{T_{1\text{yr}}} \frac{2a^{7/2} (1-e^2)^2}{3R_E^2 J_2 \sqrt{\mu_E}} \right]$$

$$= 1.718 = 98.5 \text{ deg}$$

$\uparrow$   
 $e=0, T_{1\text{yr}} = 86164.0365.25 \text{ sec}$



$G_0 =$  Greenwich meridian at  $t_0$   
 $D_0 =$  descending node at  $t_0$   
 $A_0 =$  ascending node at  $t_0$   
 ( $\lambda_{g0} = 15 \text{ deg}, \theta_{g0} = 45 \text{ deg}$ )

From the figure  $\Omega = \theta_{g0} + \lambda_{g0} + \pi = 240 \text{ deg}$

(b) Ascending node is flown after half period, i.e.:

$\Delta t = 50 \text{ min}$

$\theta_{gf} = \theta_{g0} + \omega_e \Delta t$ , then  $\lambda_{gf} = \Omega - \theta_{gf} = -3.099 =$   
 $= -177.5 \text{ deg}$

(c)  $x_0 = 0 \text{ km}$       $y_0 = 8 \text{ km}$

$\dot{x}_0 = 2 \frac{\text{m}}{\text{sec}}$       $\dot{y}_0 = 0 \frac{\text{m}}{\text{sec}}$

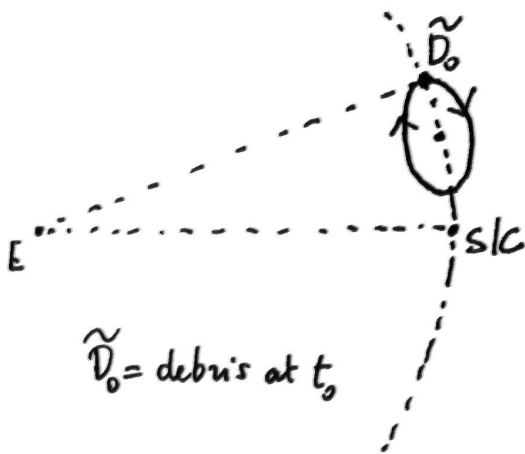
$$\begin{cases} x(t) = \frac{\sin(\omega_R t)}{\omega_R} \dot{x}_0 \\ y(t) = y_0 + \frac{2c_1}{\omega_R} [\cos(\omega_R t) - 1] \end{cases}$$

in this case, where

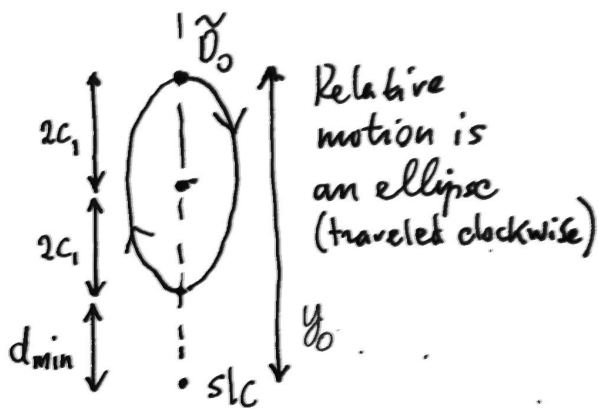
$$c_1 = \frac{|\dot{x}_0|}{\omega_R} = \frac{2 \cdot 10^{-3} \text{ km}}{2\pi} T =$$

$$= 1.910 \text{ km}$$

$$\rightarrow d_{\min} = y_0 - 4c_1 = 361 \text{ m}$$



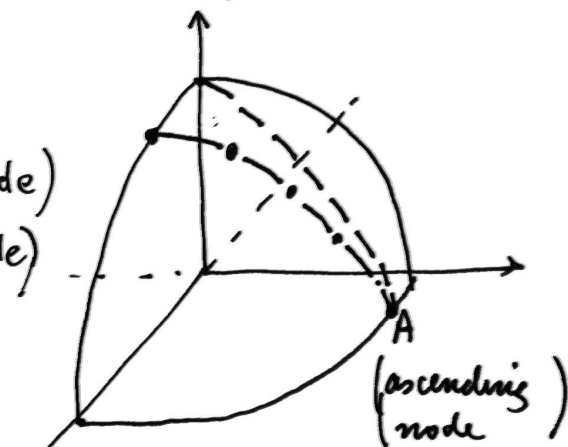
$\tilde{D}_0 = \text{debris at } t_0$



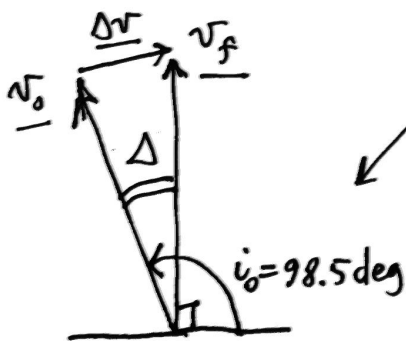
(d) Polar orbit is obtained

by a single velocity change at either A (ascending node)

or D (descending node)



$$\begin{aligned} |\Delta v| &= 2v_0 \sin \frac{\Delta}{2} \\ &= 1.102 \frac{\text{km}}{\text{sec}} \end{aligned}$$



$$\begin{aligned} \Delta &= |i_0 - i_f| \\ &= 8.5 \text{ deg} \end{aligned}$$

At A the polar orbit (curve  $\cdots$ ) intersects the sun-synchronous orbit (curve  $\cdots$ )

The same occurs at the descending node (not shown)

(e)

Initial geo long was found at point (a)

Geolong at ascending node was found at point (b)

Ground track of polar circular orbit includes straight lines traveled westward

180-deg jumps occur at polar latitude ( $L = \pm \frac{\pi}{2}$ )

The figure depicts the ground track in 1 orbital period.

(f) Given  $\rho_{800} = \rho(800 \text{ km})$  and  $\rho_{700} = \rho(700 \text{ km})$

$$\rho = \rho_0 \exp\left[-\frac{H-H_0}{H_{sc}}\right] \rightarrow H_{sc} = -\frac{H_f - H_0}{\ln \frac{\rho_f}{\rho_0}} = 88.7 \text{ km}$$

$$\rho_0 = \rho_{700}, \rho_f = \rho_{800}$$

$$H_f = 800 \text{ km}, H_0 = 700 \text{ km}$$

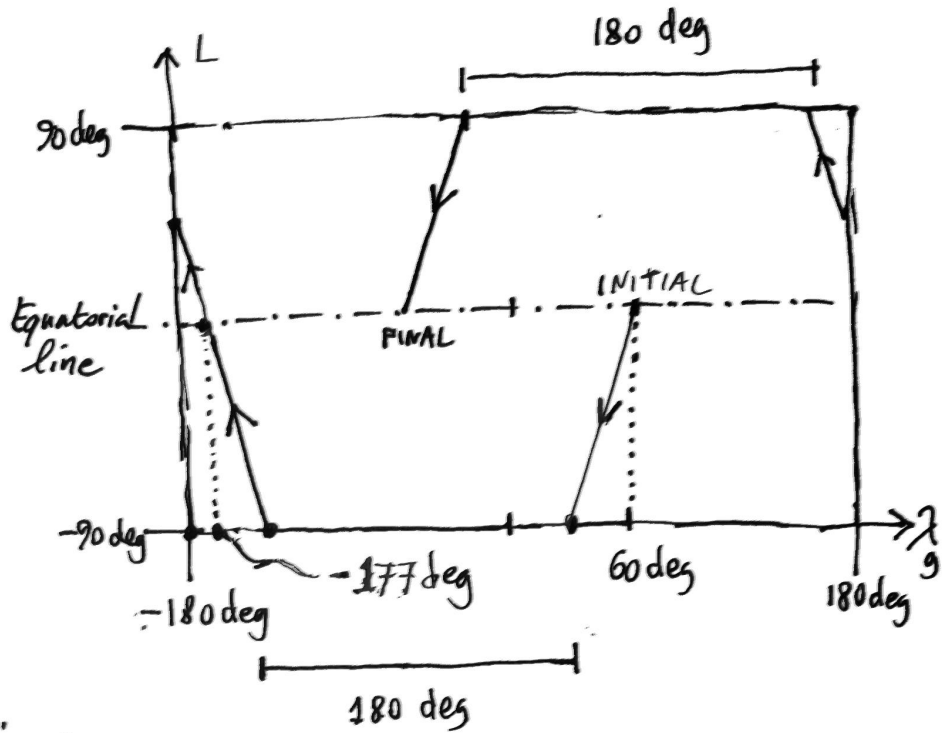
$$\rho_s = \rho_0 \exp\left[-\frac{H_s - H_0}{H_{sc}}\right]$$

$$= 1.601 \cdot 10^{-5} \frac{\text{kg}}{\text{km}^3}$$

$$\rightarrow a_{fin} = \left[ a_{ini}^{\frac{1}{2}} - \frac{G_0 S}{m} \rho \sqrt{\mu_E} \Delta t \right] = 7135.860 \text{ km}$$

$$\rightarrow \Delta r = 773 \text{ m} \quad \left( \begin{array}{l} \text{decrease in} \\ \text{altitude} \end{array} \right)$$

$$\Delta t = 90.86400 \text{ sec}$$



③ See solution of exercise 3, set 7